

A photograph of Saturn and its rings, centered in the frame. The planet is a pale yellowish-white, and the rings are a mix of grey and white, showing some darker gaps. The background is black.

The Smallest Free Particles in Saturn's Rings

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Acknowledgements

- Philip Nicholson (Cornell), Matthew Hedman (University of Idaho) - Co-authors
- Joseph Burns, Rachel Bean, Richard Lovelace (Cornell) - Dissertation Committee

Outline



1. Introduction

- Saturn's Rings & Cassini
- Occultations
- Particle Size Distributions

2. Solar Occultations

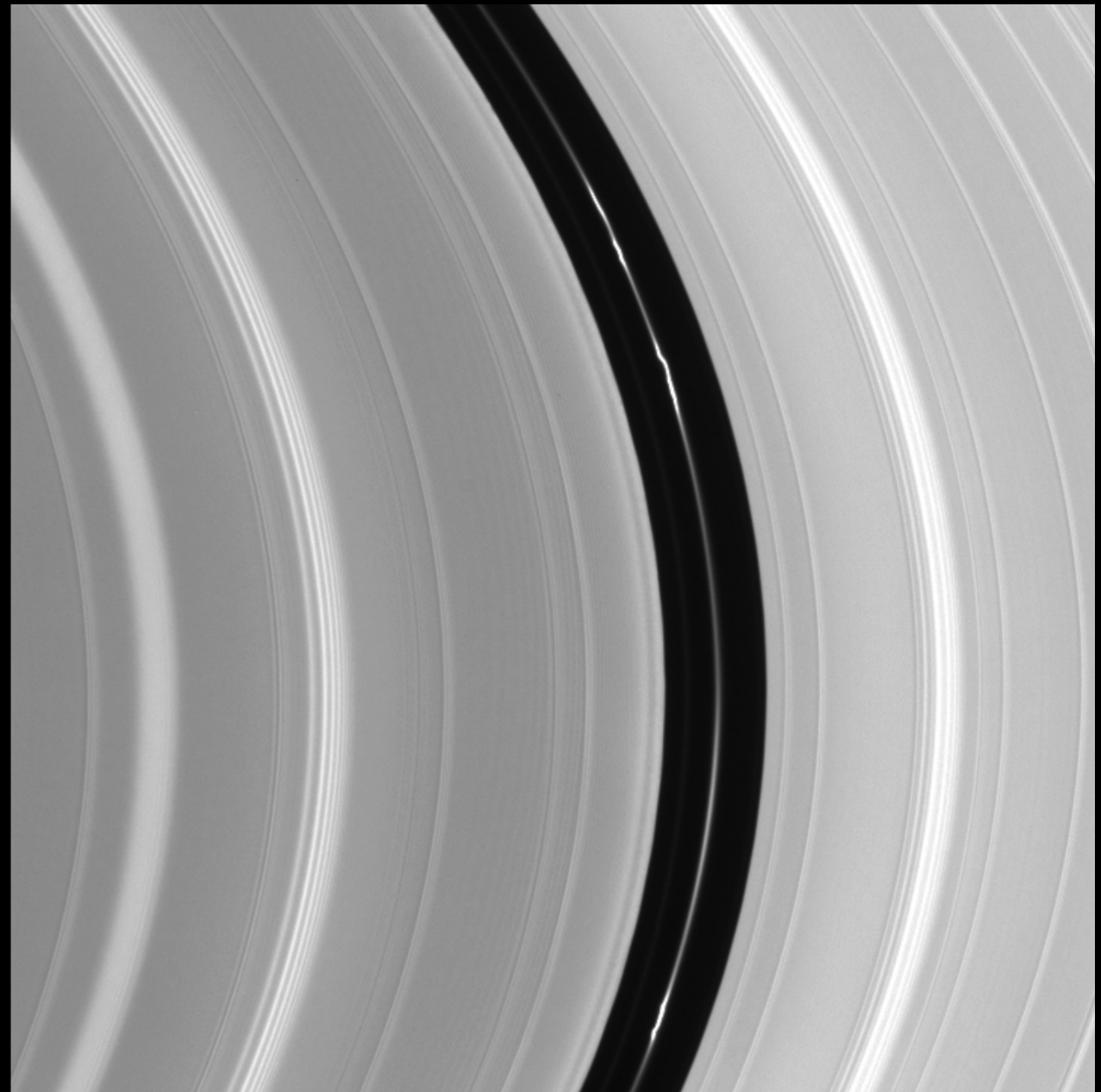
- Data Analysis
- The C Ring
- The A Ring (is more complicated)

3. Conclusions about Ring Particle Size Distribution

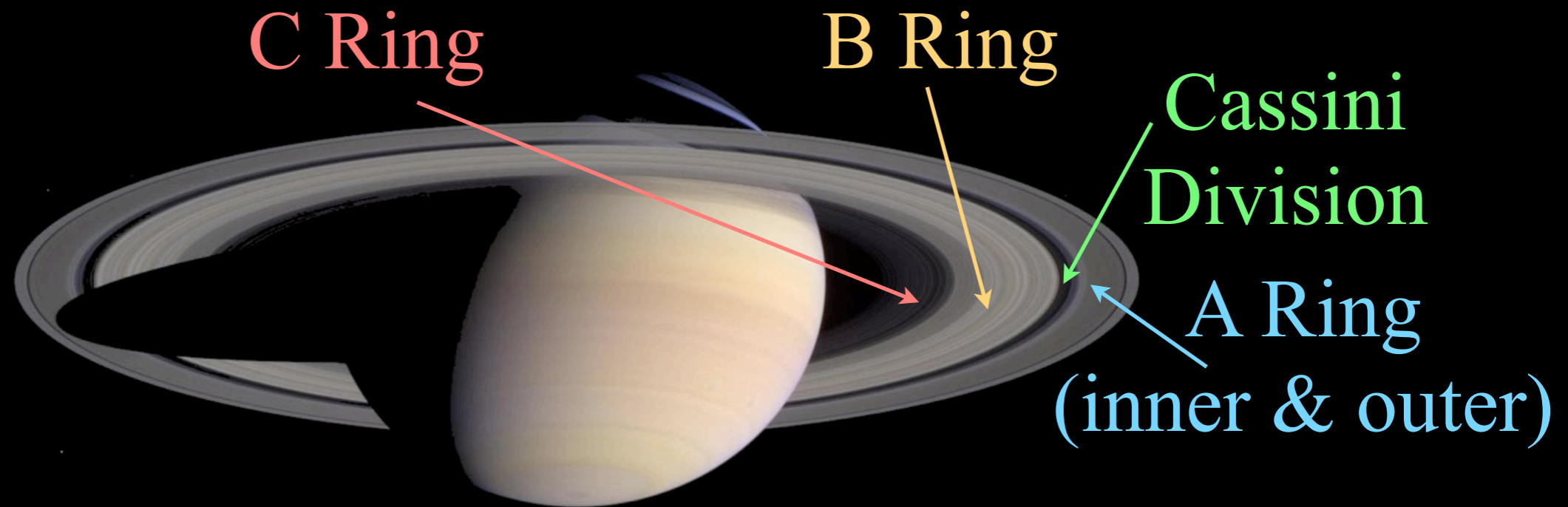
Motivation

Why do we care about what size Saturnian ring particles are?

- Better models of Saturn ring dynamics.
- A better understanding of Saturn's rings and their possible origin and evolution.
- Rings are our local debris disks (common objects in the Universe).



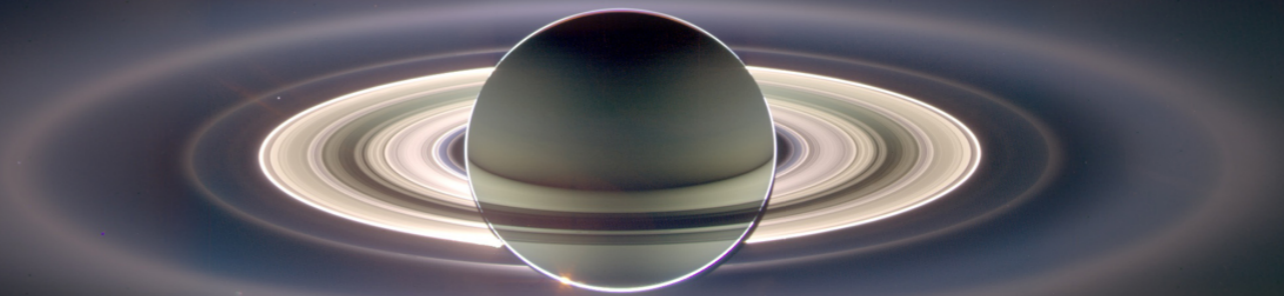
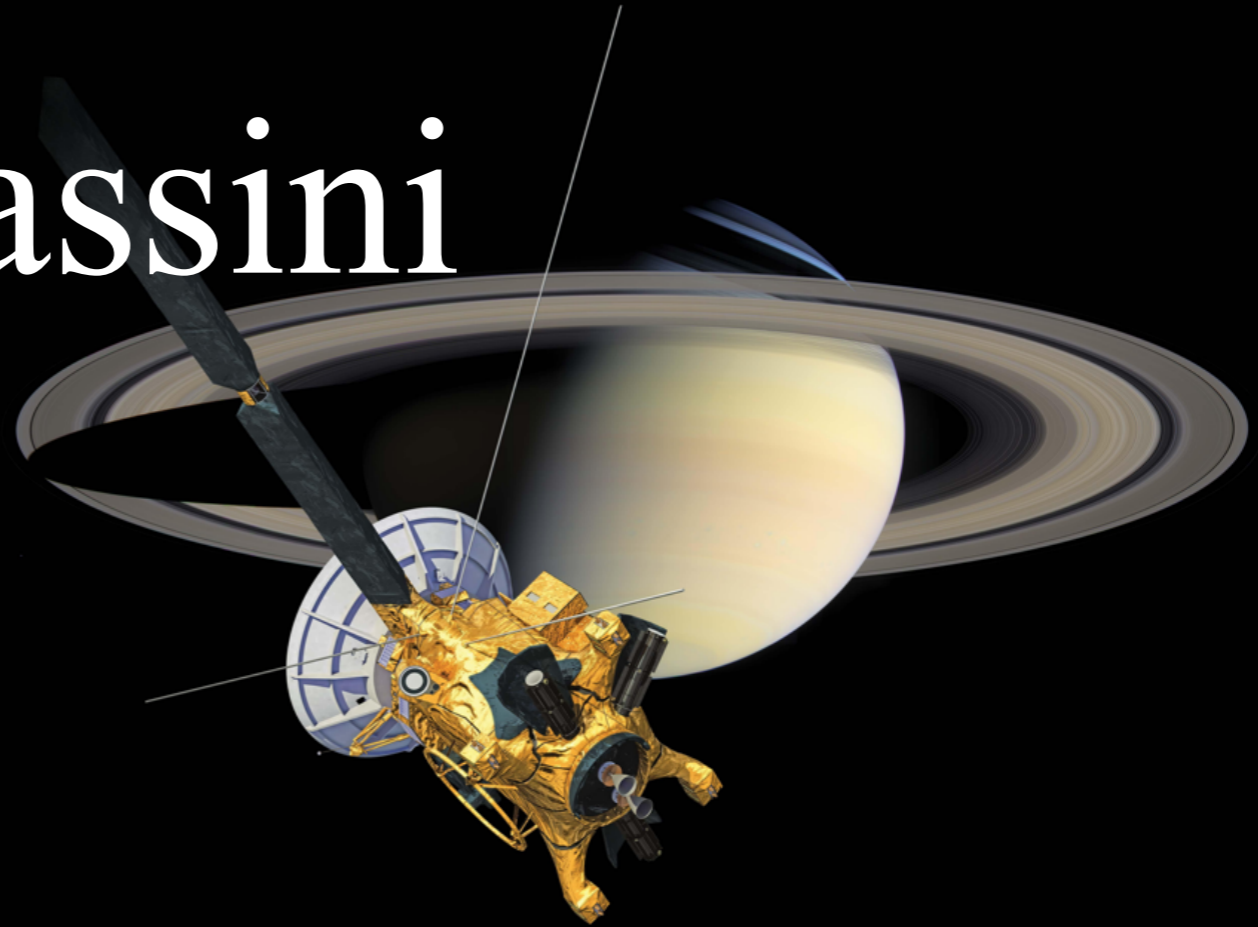
The Rings of Saturn



- The Main Rings are made of solid pieces of water ice + trace impurities.
- Rings exist because tides $>$ self-gravity.
- The rings have distinct regions with different properties.

Cassini

- Cassini entered orbit around Saturn in 2004.
- Constantly changing geometry as repeated close encounters with Titan alter the orbit.
- Able to achieve resolutions and geometries impossible from Earth



Visible-Infrared Mapping Spectrometer

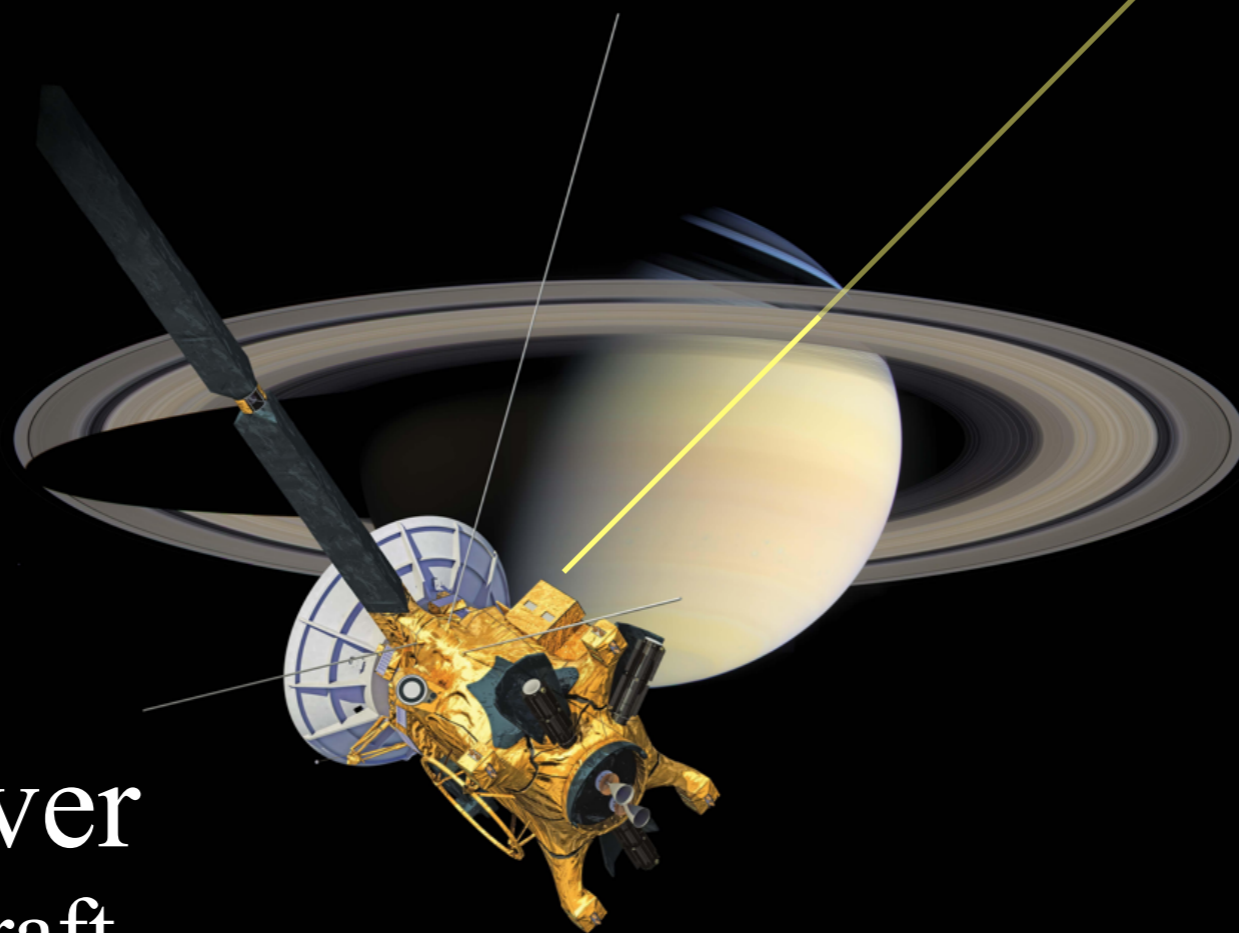
- 256 infrared channels.
- Wavelength range from 0.9 to 5.2 μm in occultation mode.
- Pixel size 0.5 x 0.5 milliradians
- Boresight allows solar observations
- Capable of imaging pixel-by-pixel.



VIMS Occultations

Source
Sun or star

Receiver
Spacecraft

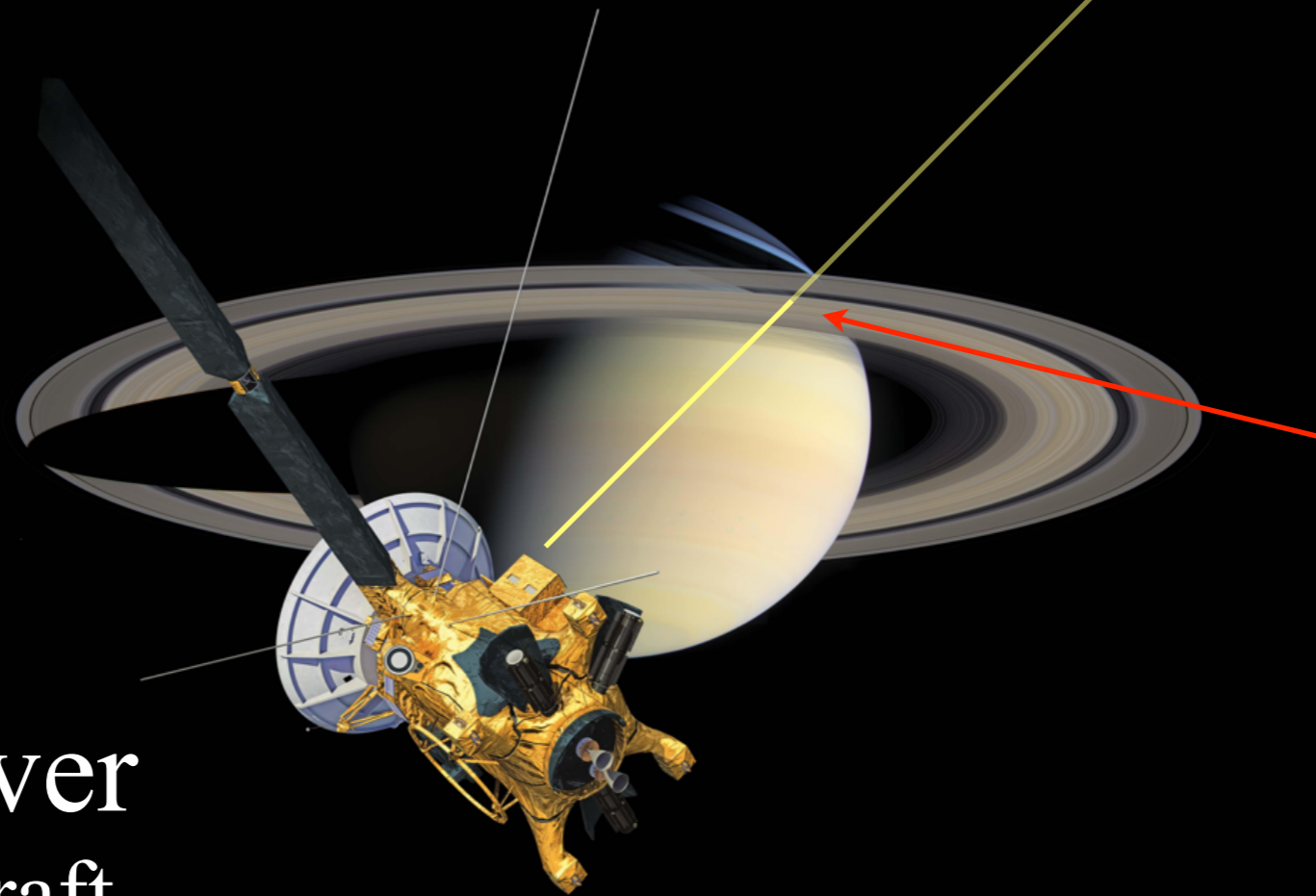


VIMS Occultations

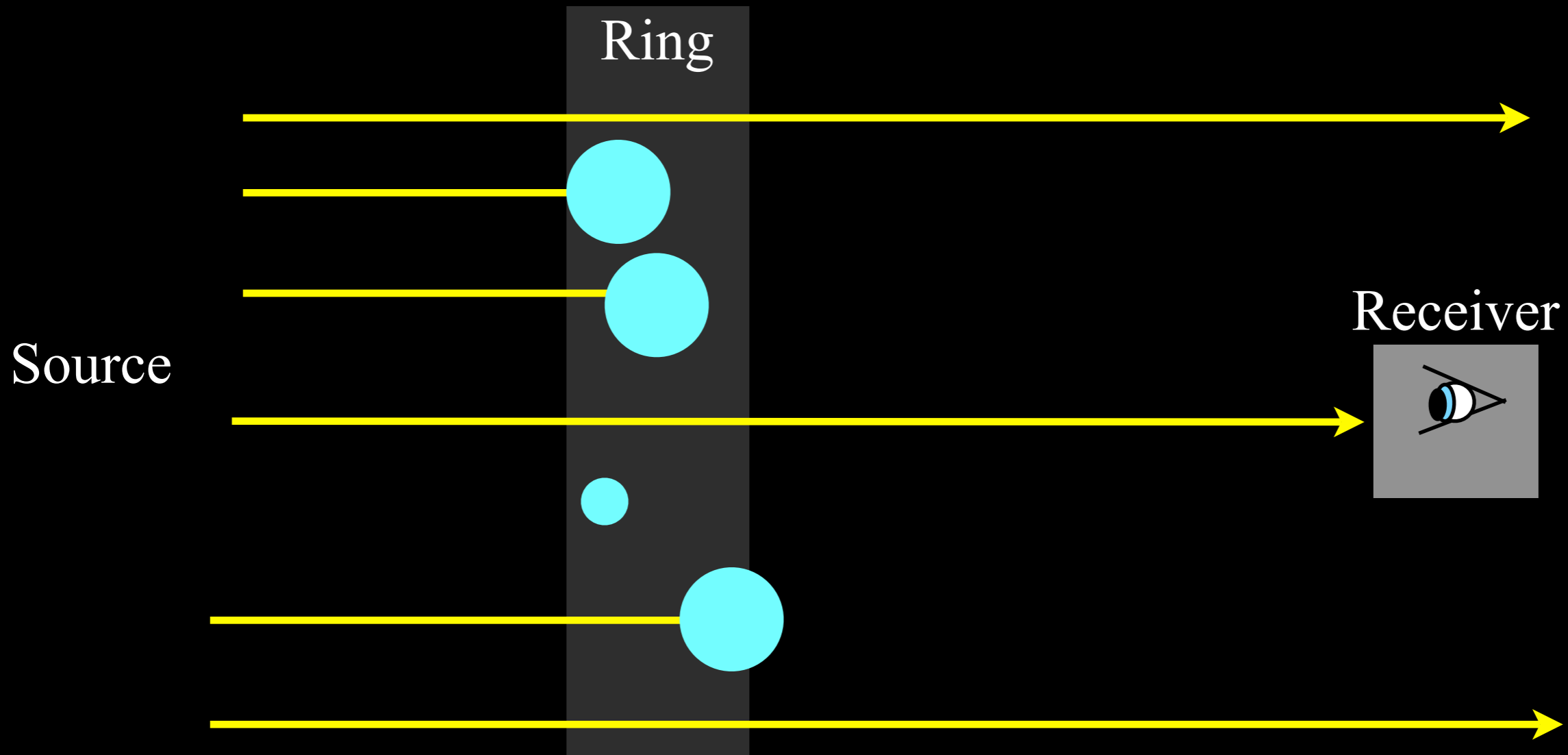
Source
Sun or star

Receiver
Spacecraft

Target
Rings

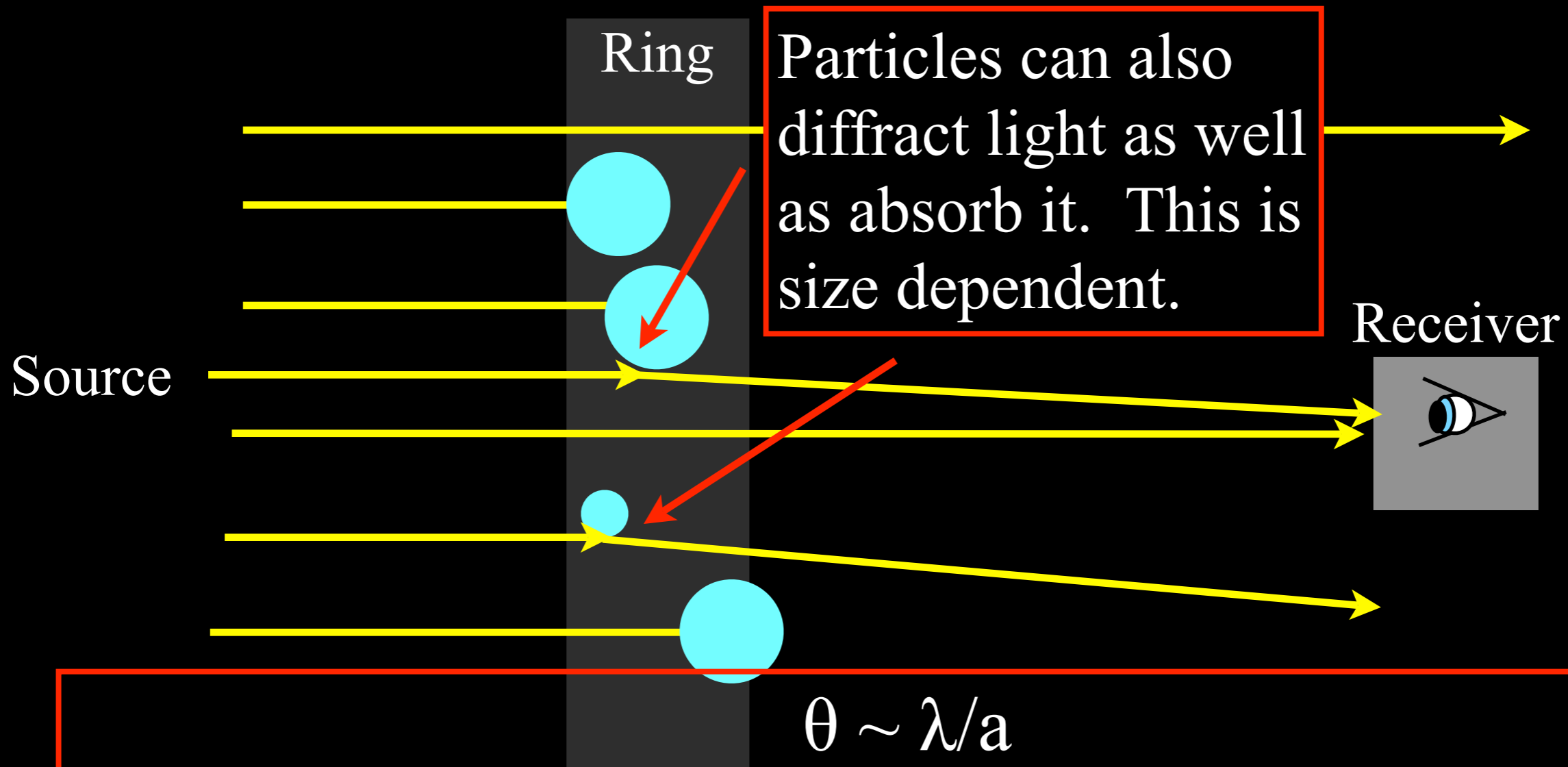


Occultations



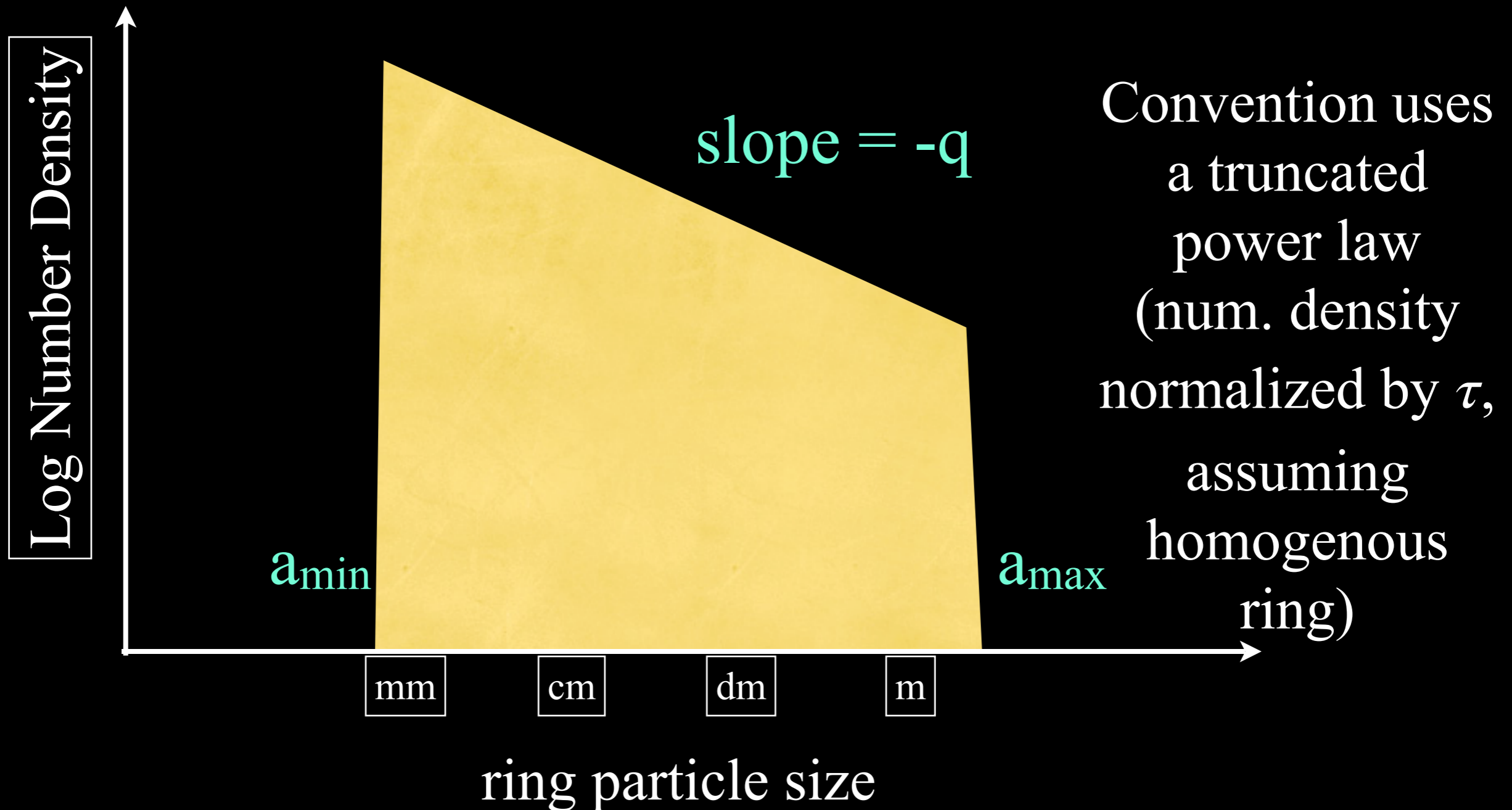
Some light makes it through the ring,
Some light is absorbed /reflected by ring particles.

Occultations



0.5 mrad @ 2 μm \longrightarrow mm-sized ring particles
Diffracted light would make an aureole around the source.

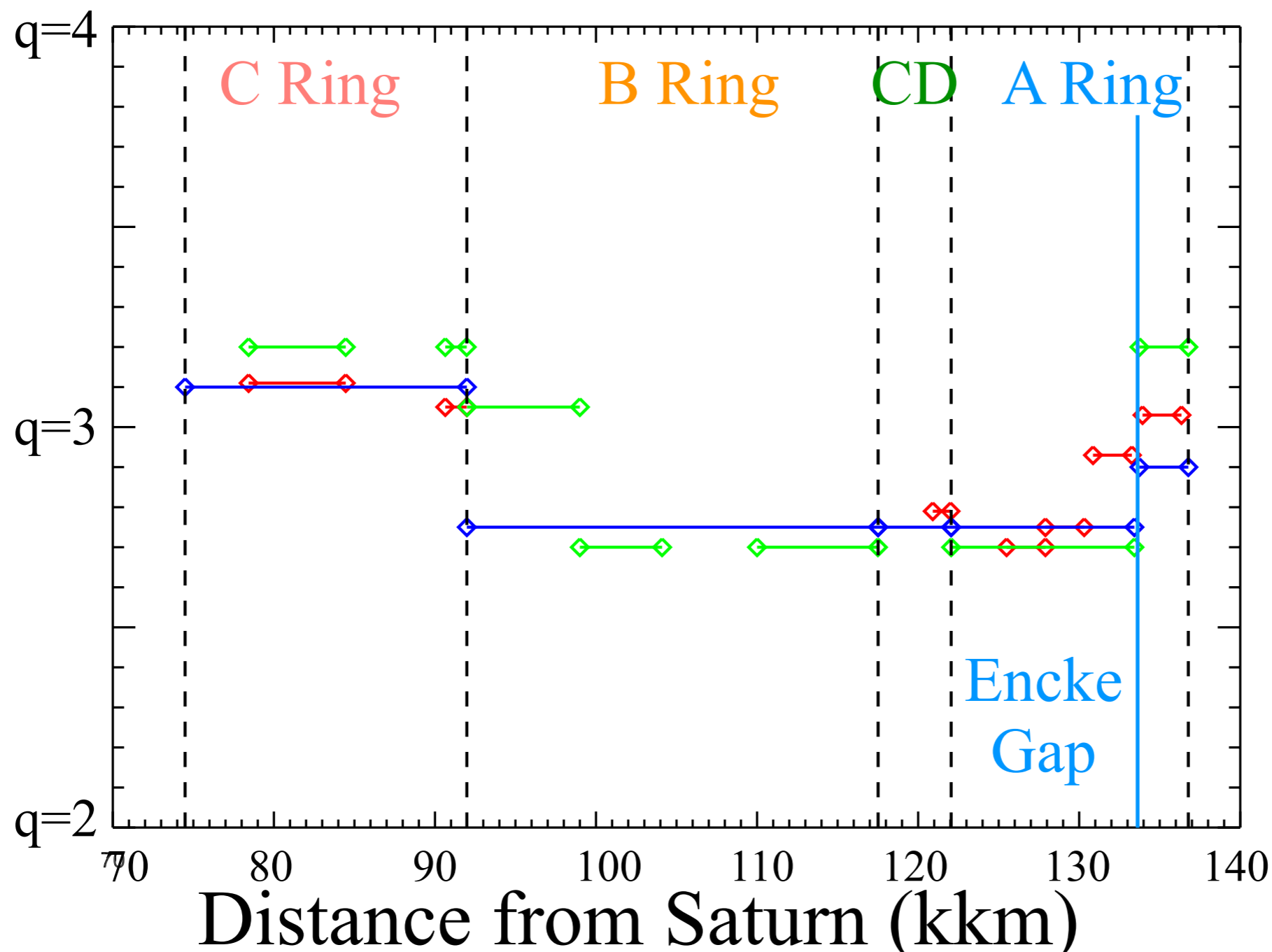
Modeling the Particle-Size Distribution



Previous Knowledge: General

- There is little free ring material smaller than millimeter-sized ring particles.
- Not much is known about the exact lowest cutoff size. (a_{\min})
- The size-distribution steepens markedly at ~ 5 -10 meters, acting like an upper cutoff (a_{\max})
- The size distribution of the ring particles between centimeters and meters acts like a power law of $q \sim 3$.

Previous Knowledge: Ring Regions

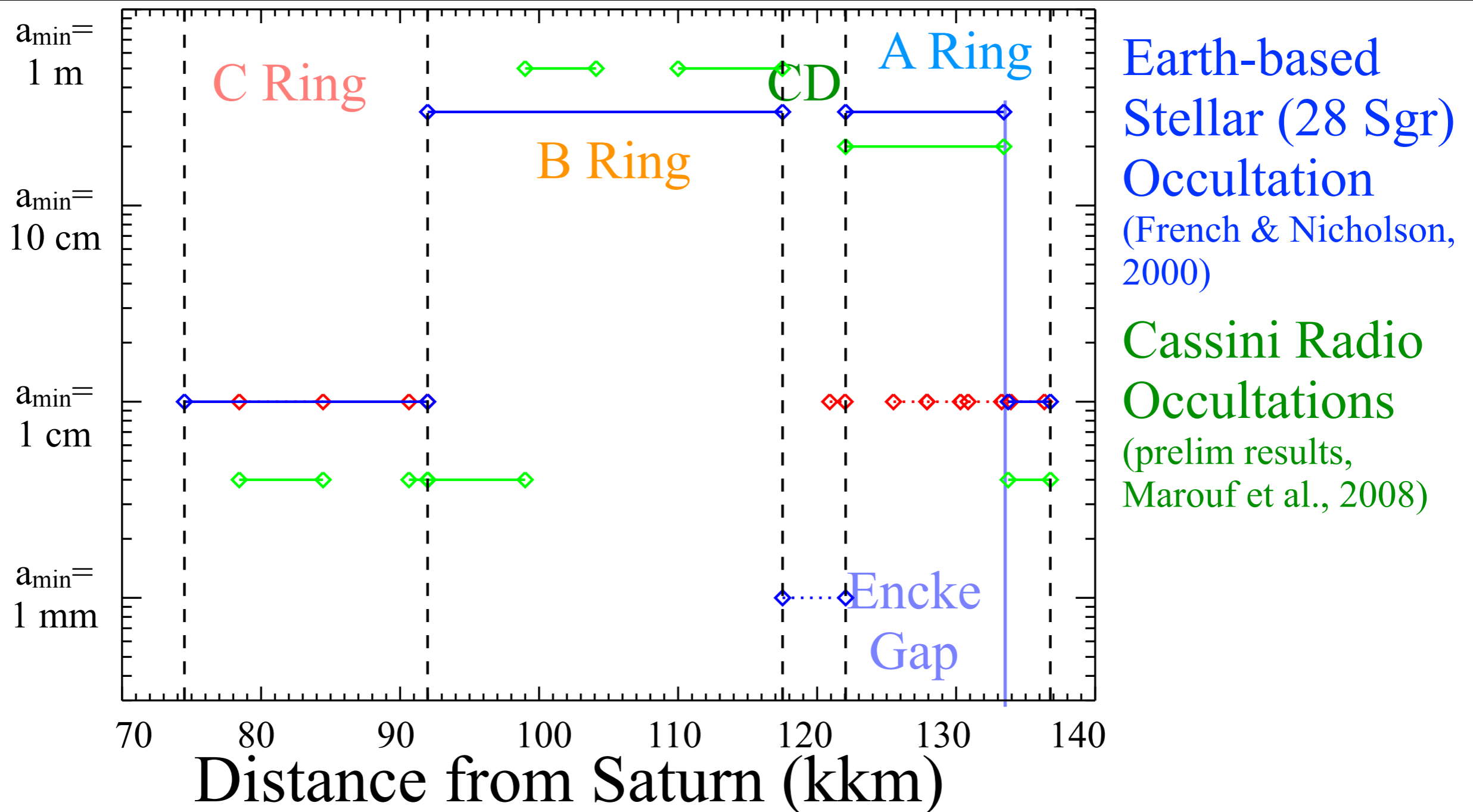


Voyager Radio
Occultation
(Marouf et al., 1983,
Zebker et al., 1985)

Earth-based
Stellar (28 Sgr)
Occultation
(French & Nicholson,
2000)

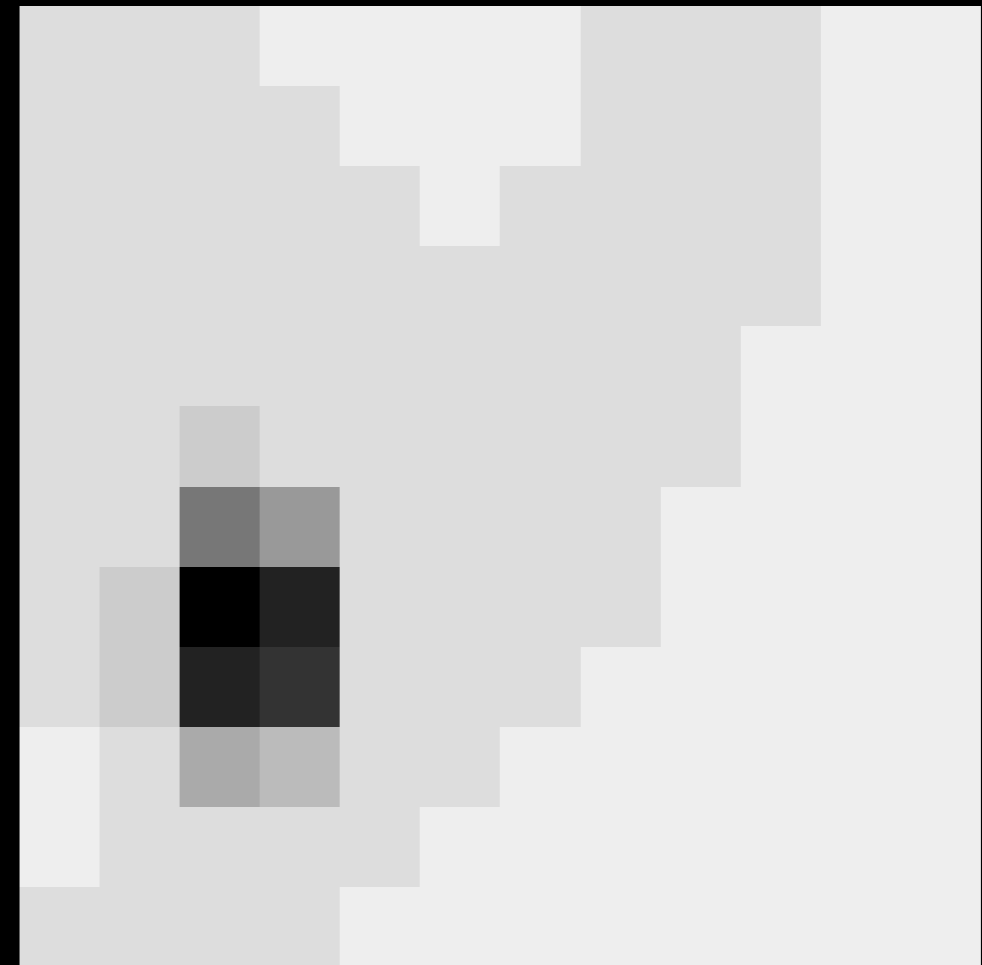
Cassini Radio
Occultations
(prelim results,
Marouf et al., 2008)

Previous Knowledge: Ring Regions



Solar Occultations

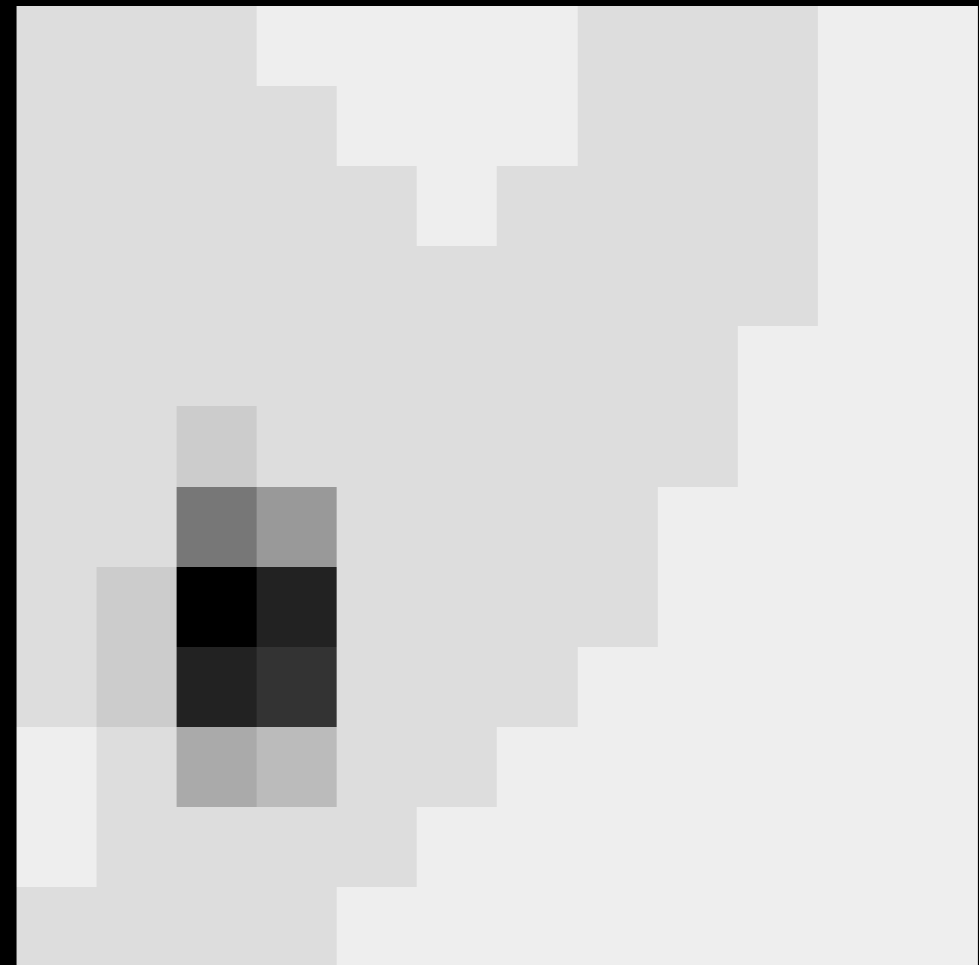
- Details in Harbison et al., 2013, in *Icarus*
- Taken as 12 x 12 pixel images.
- 6 occultations had sufficient S/N.



Reverse-contrast
VIMS image of the Sun

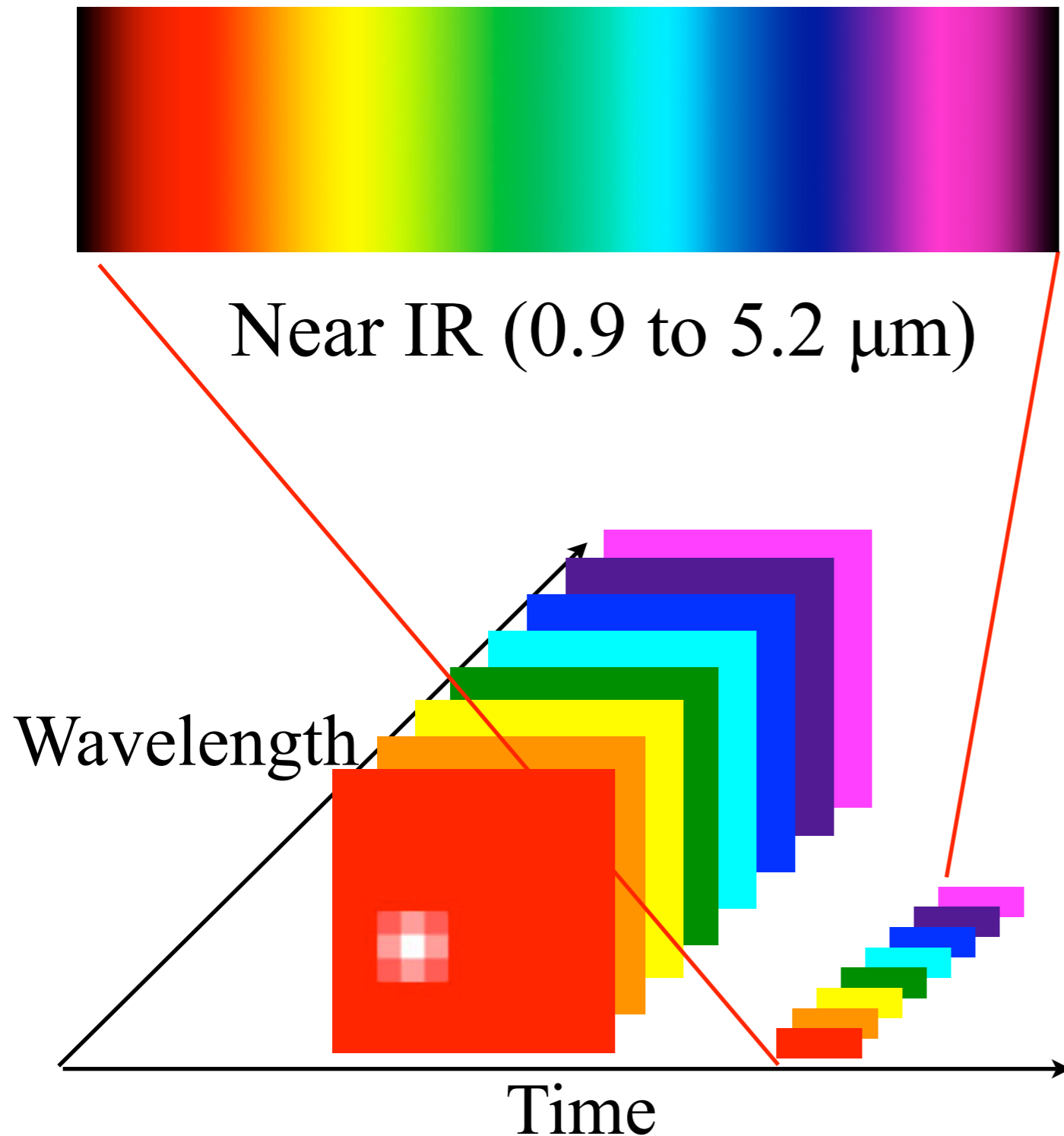
Instrumental Effects

- This is an image of the Sun outside the rings.
- Instrument-scattered light is stronger than the halo from ring particle diffraction.
- How to measure it?



Background is $\sim 1/10$ of
peak solar signal

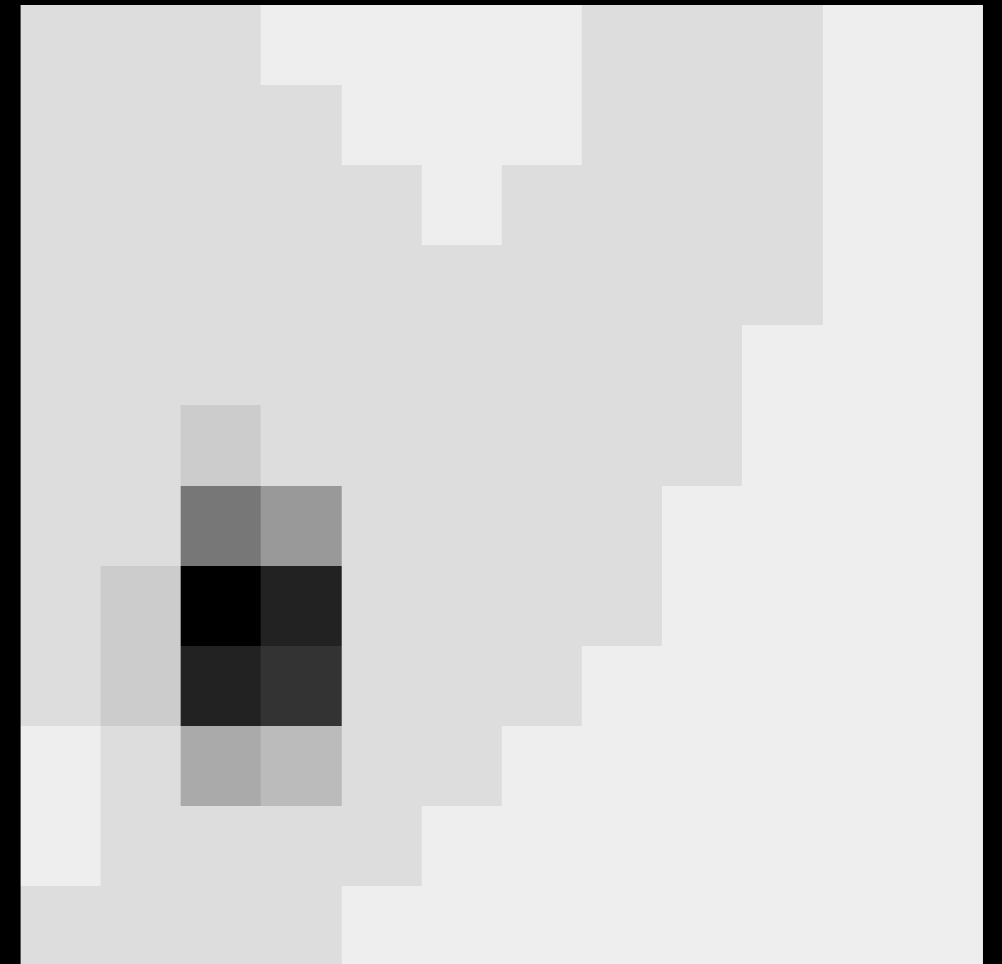
VIMS Imaging (Theory)



- Light enters the solar port and is attenuated.
- A small mirror directs part of the field of view into the spectrometer.
- This creates a spectrum of one pixel.
- The mirror tilts to direct the next pixel's light onto the spectrometer.

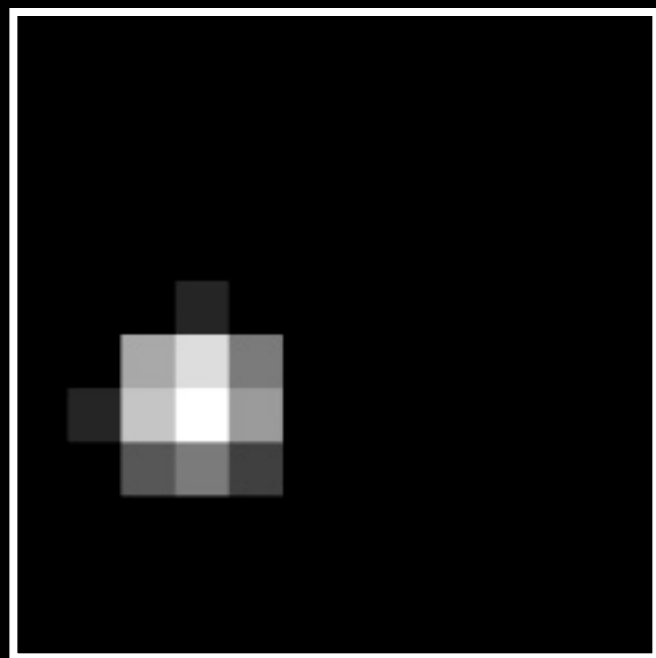
VIMS Imaging (Practice with Solar Port)

- Light enters the solar port and is attenuated and scattered.
- A small mirror directs part of the FOV into the spectrometer.
- Scattered light from the entire FOV also enters the spectrometer.
- The mirror tilts to direct the next pixel's light onto the spectrometer. The scattered light entering the spectrometer does not change (as much).



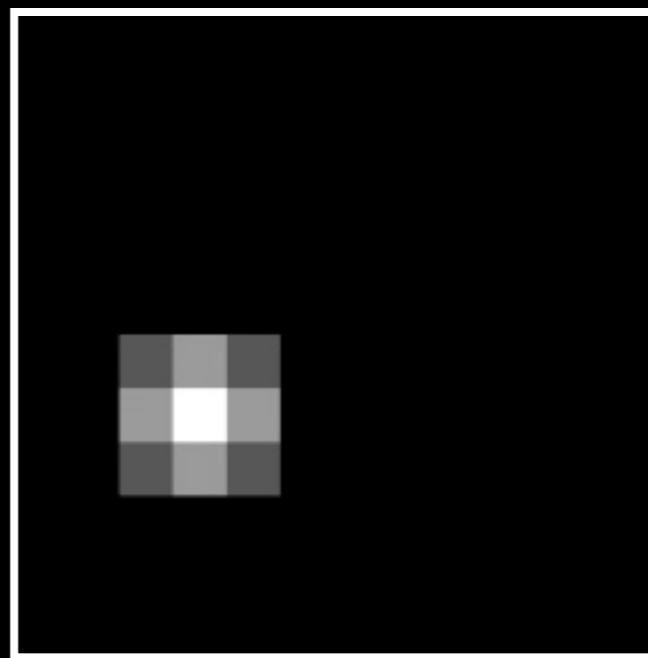
Measuring Diffraction

(Simulated images)



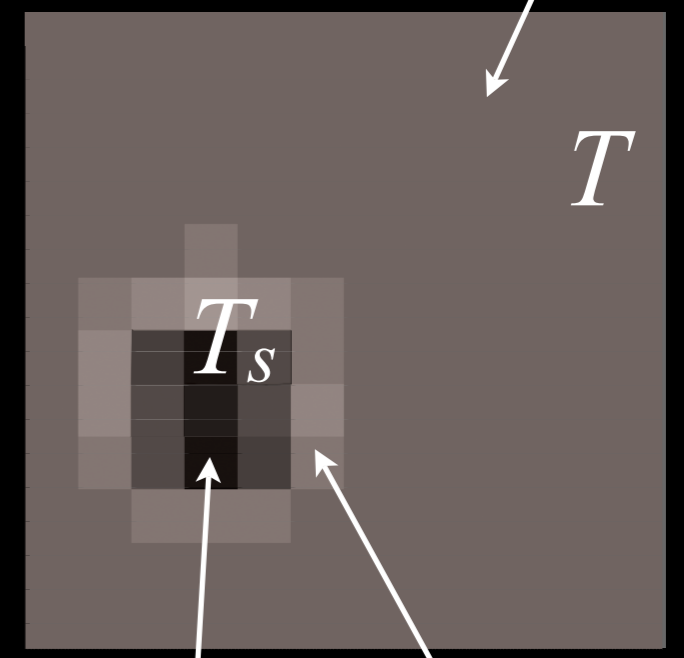
Ring Cube

\div



Mean of
Empty Cubes

$=$

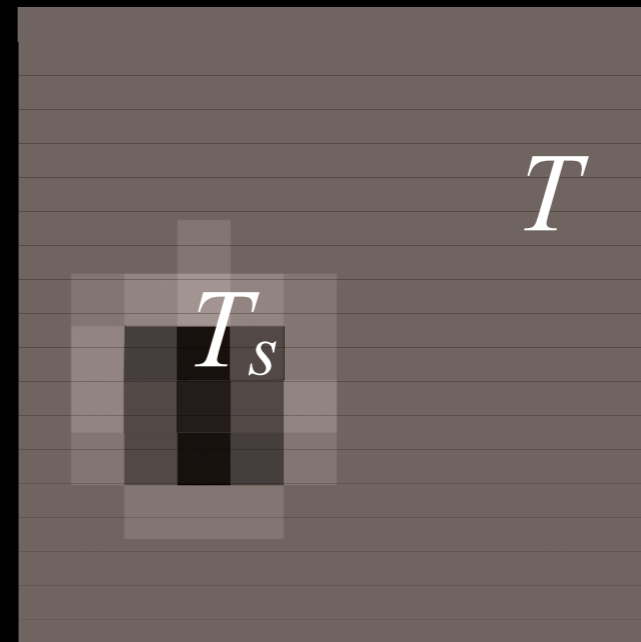


Sun's image is
dimmer
Surrounding pixels
are brighter.

BG large enough to
work as reference.

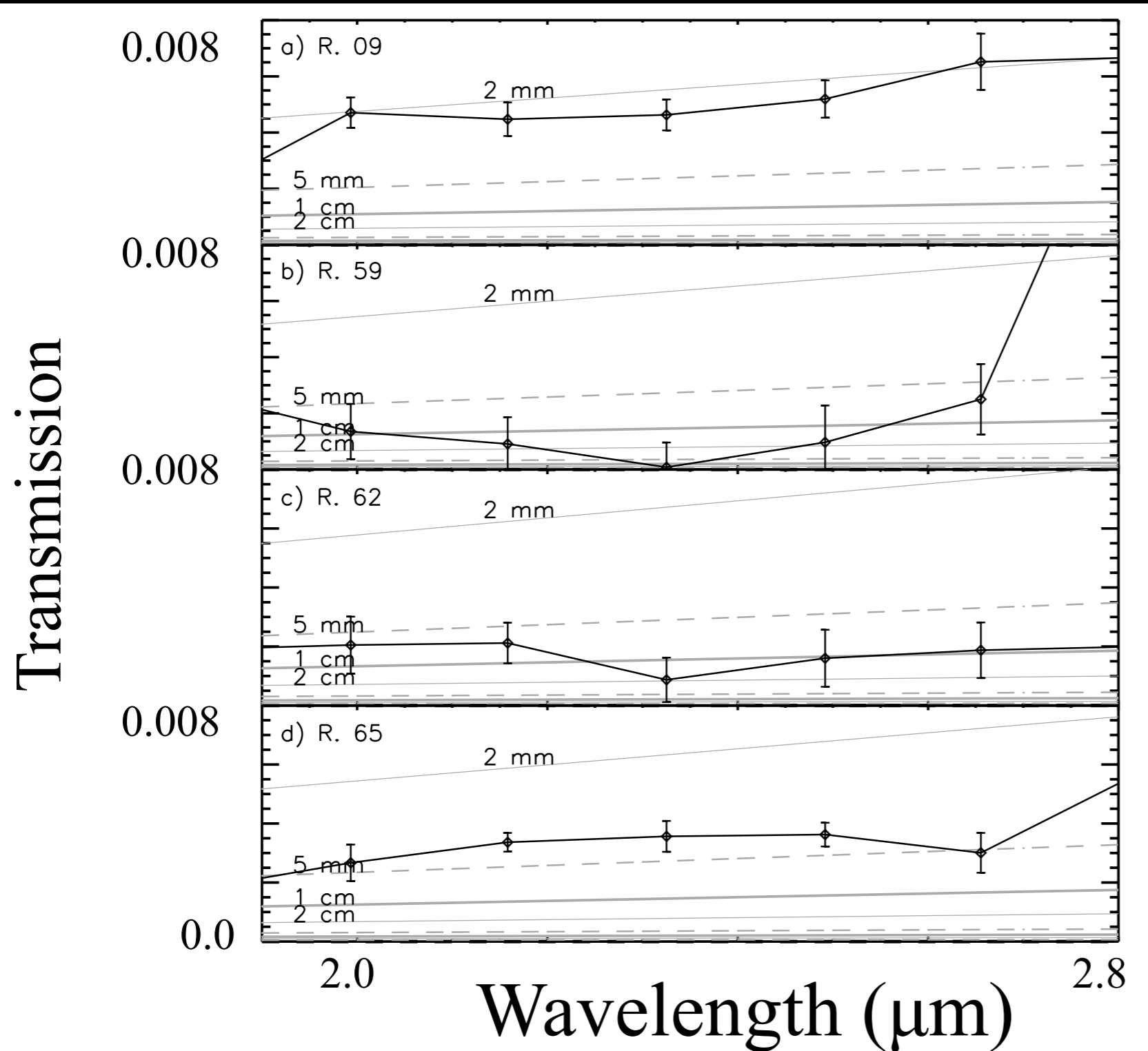
Measuring Diffraction

- Transmission found by fitting Gaussian function to sun + background.
- Difference in transmission \Rightarrow fraction of light diffracted to θ greater than 0.5 mrad \Rightarrow measure of particle properties.



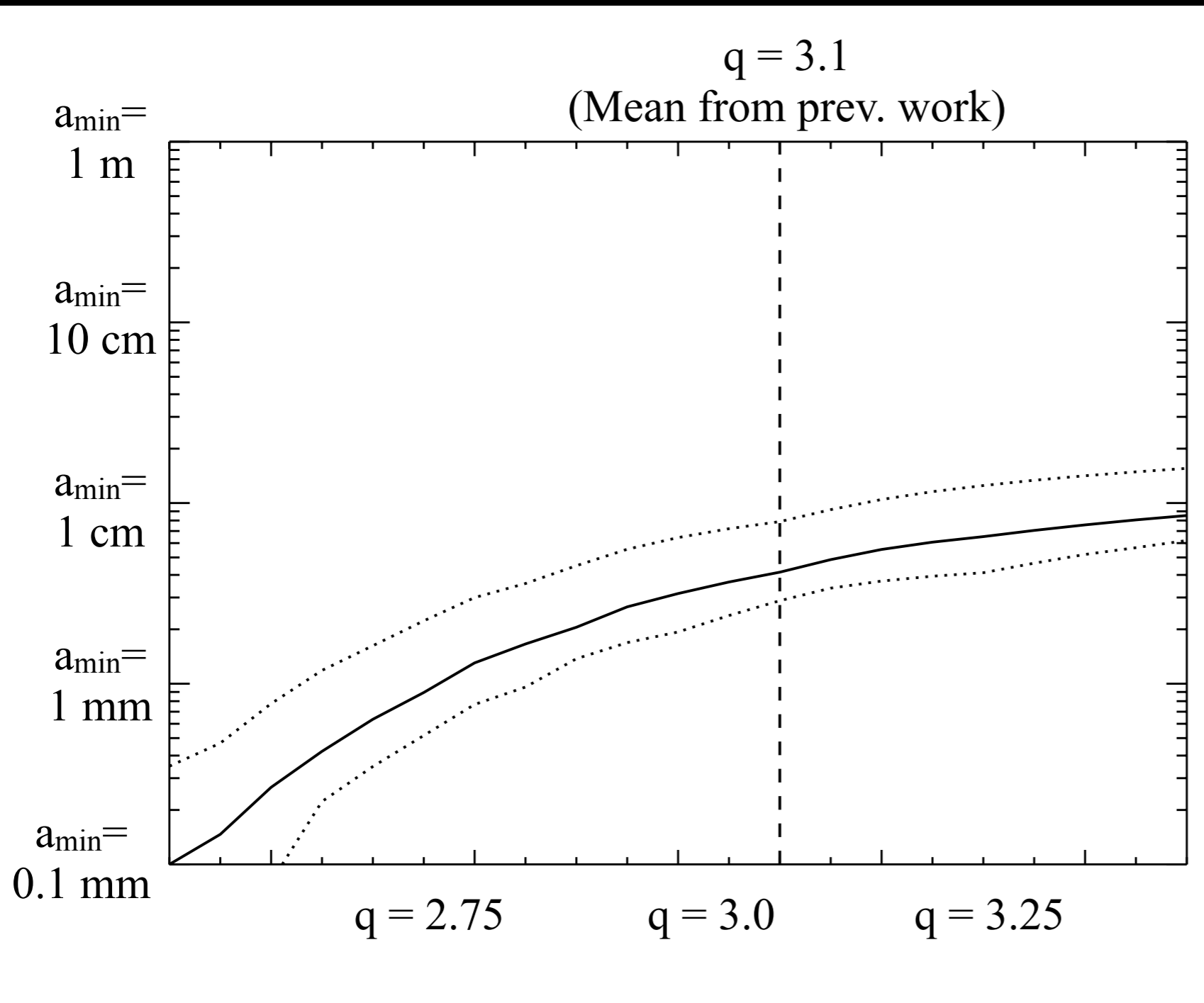
Results

The C Ring



Signal is an average over many cubes and binned in wavelength. 3 positive detections, 1 partial positive, 1 negative detection.

Minimum Particle Size in the C Ring



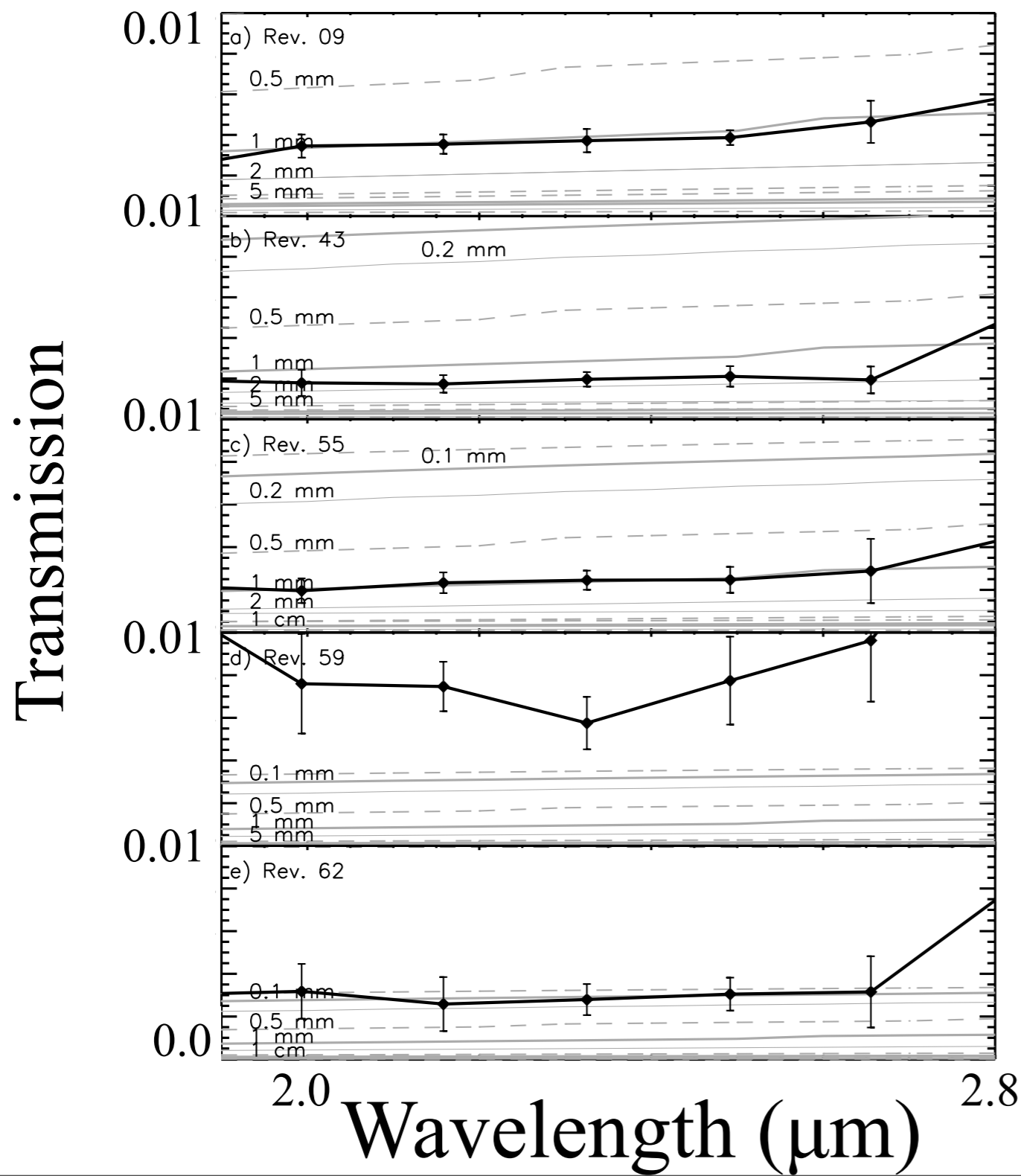
Using $q = 3.1$,
this work finds

$$a_{\min} = 4.1^{+3.8}_{-1.3} \text{ mm}$$

Comparison: C Ring

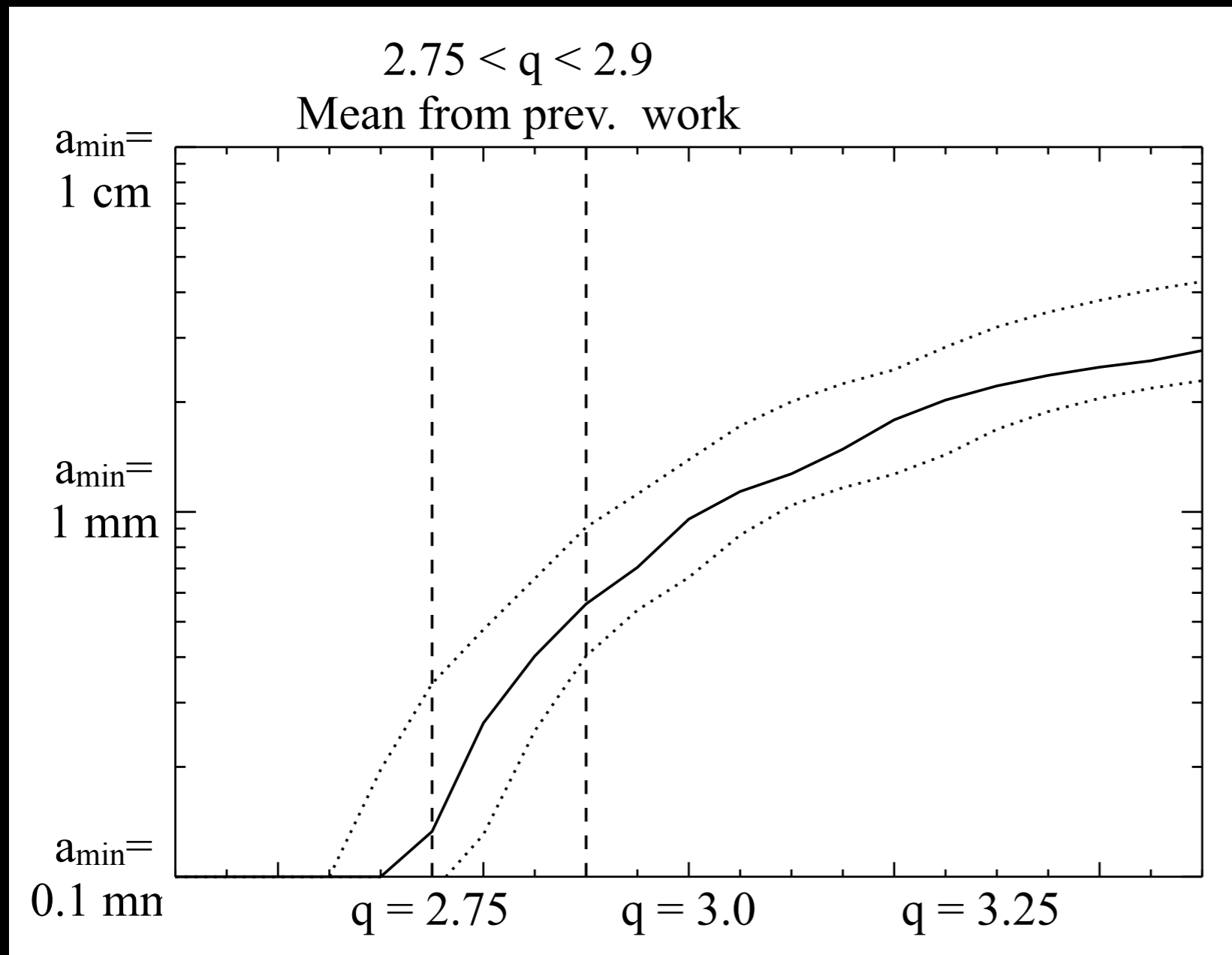
- 28 Sgr \sim 1 cm in C Ring (French & Nicholson, 2000)
- Cassini RSS = 4 mm in C Ring (Marouf, 2008)
- **This work = 4 mm in C Ring for $q = 3.1$**

The A Ring



Signal is an average
over many cubes
and binned in
wavelength.
5 positive
detections, 1
negative detection.

Minimum Particle Size in the A Ring



Using $q = 2.9$,
this work finds

$$a_{\min} = 0.6^{+0.4}_{-0.2} \text{ mm}$$

Using $q = 2.75$,
this work finds

$$a_{\min} < 0.3 \text{ mm}$$

Comparison with Previous Results

- 28 Sgr = 30 cm in inner-to-mid A Ring (French & Nicholson, 2000)
- Cassini RSS = 20 cm in inner-to-mid A Ring (Marouf, 2008)
- **This work = 0.6 mm (*or less!*) in inner-to-mid A Ring**
- However, previous work does not include self-gravity wakes.

Self-Gravity Wakes

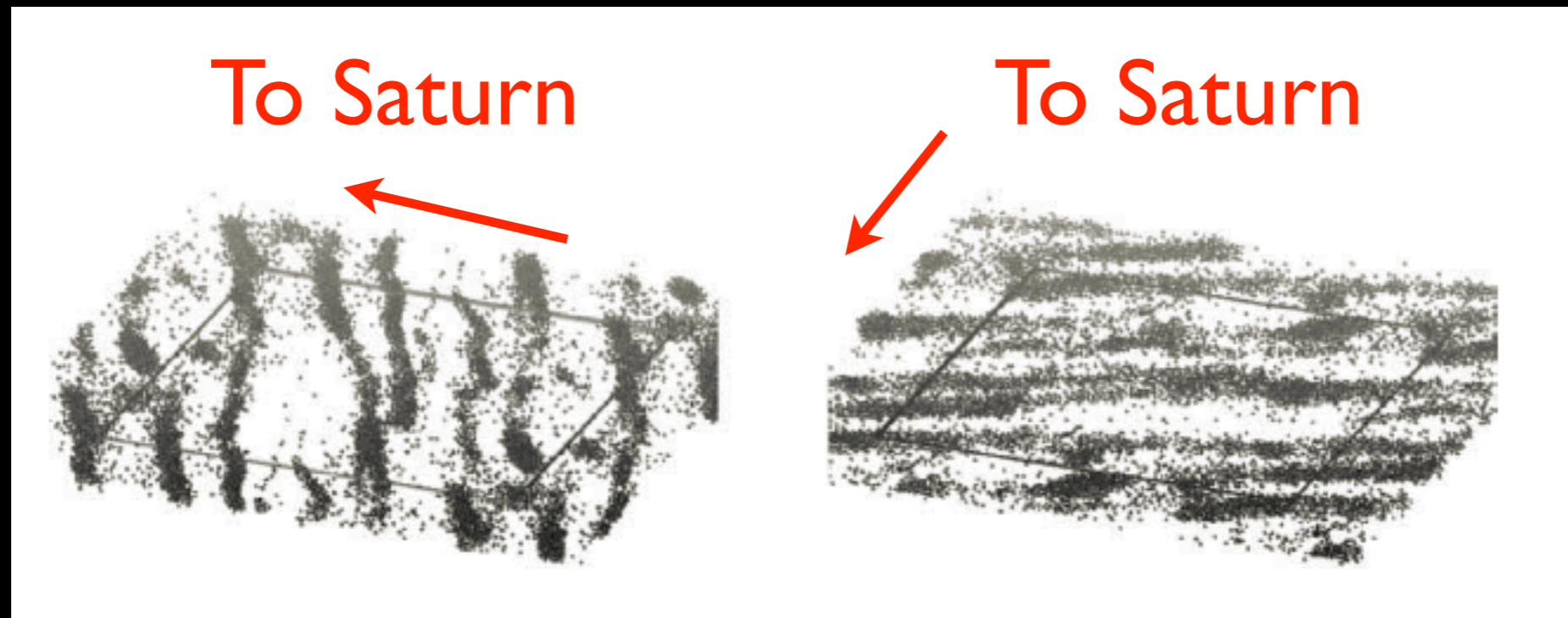


Figure from Salo et. al, 2004

- Seen seen in computational modeling of A Ring (in early 2000s).
- Explains azimuthal asymmetries seen in the A Ring pictures & occultations.

Self-Gravity Wakes

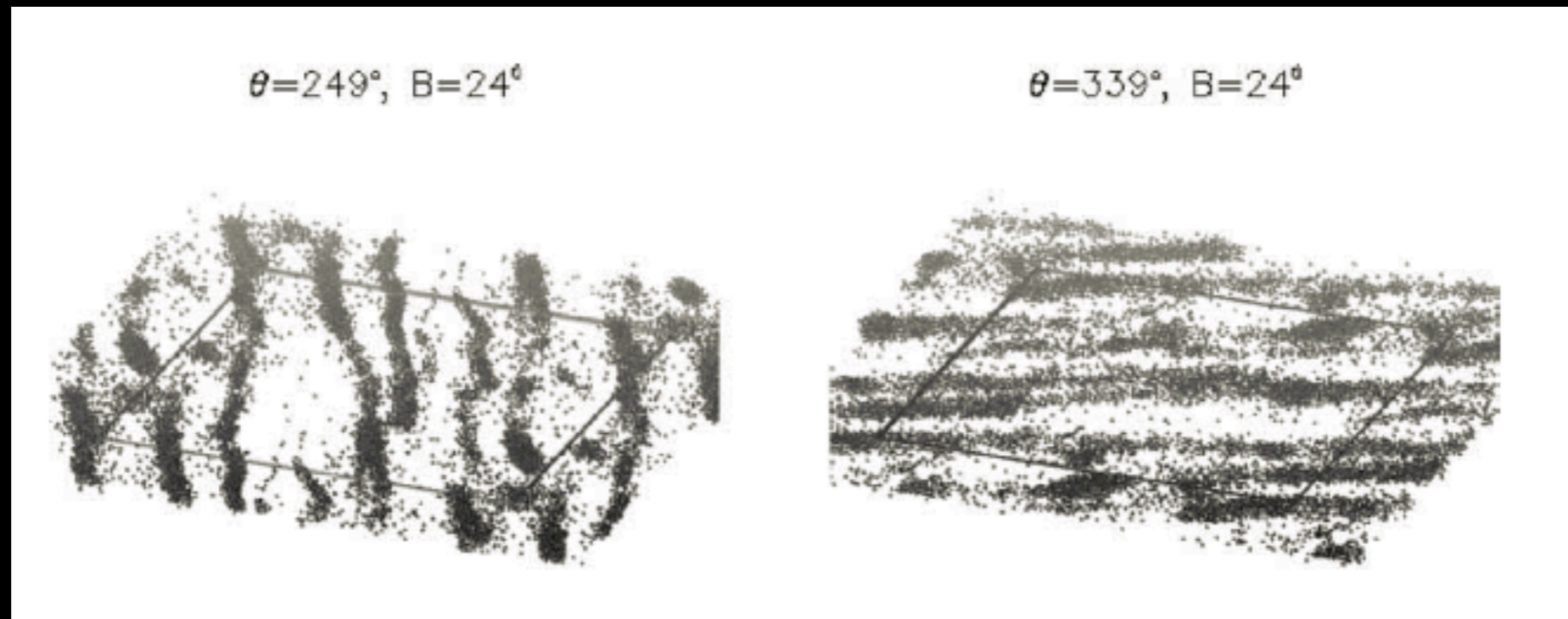


Figure from Salo et. al, 2004

Wakes are a balance between the ring's self-gravity and tides from Saturn.

Wakes are aggregates of ring particles. A wake is temporary, but the A ring *always* has self-gravity wakes

Self-Gravity Wakes

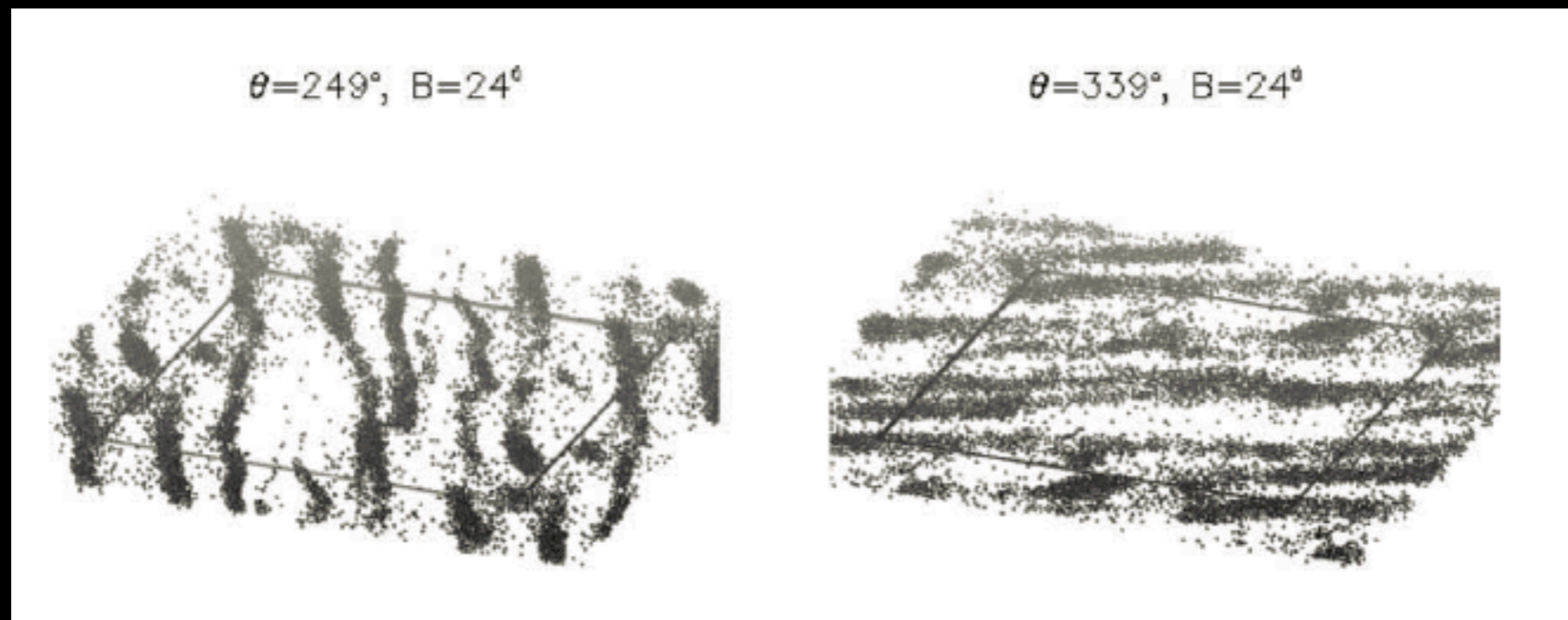


Figure from Salo et. al, 2004

- Wakes can violate a model assumption \Rightarrow that num. density comes from τ .
- However we know wake properties and can model them to get the correct num. density.

Ignoring Self-Gravity

Wakes

- Wakes mean fewer particles *of all sizes* are free to interact with light for a given τ .
- A model with wakes would have less effect from (free) small particles than one without.
- But, in many cases no wakes and few small particles produce the same effect as wakes.
- Unlike previous experiments, I am insensitive to larger particles, so my results *cannot be explained* without mm-sized particles.

Conclusions: Particle Size Distribution

- We confirm minimum particle sizes of $a_{\min} \sim 4$ mm in the **C Ring**.
- The **inner-to-mid A Ring** has minimum particle sizes of $a_{\min} < 1$ mm, over two orders of magnitude less than previous results.
- Modeling the **inner to mid A Ring** *must* take the self-gravity wakes into account.